

SO FAR:

(C6)

- state described by wavefn. $\psi(x, t)$
- time evolution of wavefn determined by Schrödinger eqn:

$$i\hbar \frac{\partial \psi}{\partial t} = -\frac{\hbar^2}{2m} \frac{\partial^2 \psi}{\partial x^2} + U(x)\psi$$

↑ potl. energy of electron
at x .

- states w. definite energy E have wavefns:

$$\psi(x, t) = \psi_E(x) e^{-i\frac{E}{\hbar}t}$$

where

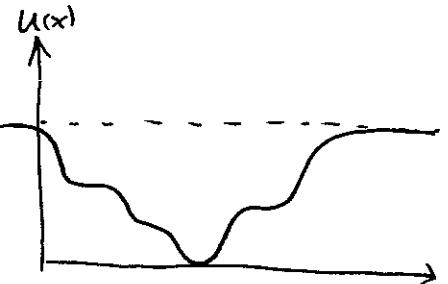
$$-\frac{\hbar^2}{2m} \frac{\partial^2 \psi_E(x)}{\partial x^2} + U(x)\psi_E(x) = E\psi_E(x)$$

TIME
INDEP.
SCHR. EQN.

RA

QM general recipe:

step 1: determine potential energy fn $U(x)$ to specify problem



step 2: use the time indep. S.E. to find energy eigenstates

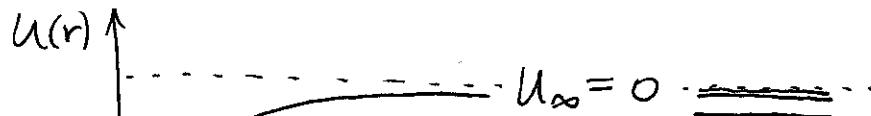
* normalizable solns with $E < U_{\infty}$ exist only for specific energies *

→ these are the wavefns for BOUND STATES.

* bound state energies determine most observable properties of system *

e.g. H atom

Bound state energies $E_n = -\frac{13.6 \text{ eV}}{n^2}$



$$E_2 = -\frac{13.6}{4} \text{ eV}$$

$$E_1 = -13.6 \text{ eV}$$

* Discrete freqs.
in atomic spectra
explained by
transitions $E_a \rightarrow E_b$

photon emitted w.
energy $\Delta E = E_a - E_b$

Orbital sim
CLICKER

* There is a state of minimum energy *

\Rightarrow Atoms are stable.

* Energy eigenstates are STATIONARY: $|\psi(x,t)|^2$ indep of time \therefore no acceleration \therefore no radiation.

~~DO SQ. WELL SIM~~ \rightarrow CLICKER.

General state: superposition of energy eigenstates

$\Rightarrow |\psi(x,t)|^2$ not indep. of time, but time dependence follows from time dep. of eigenstates

$$\psi(x,0) = \sum c_n \psi_{E_n}(x) \Rightarrow \psi(x,t) = \sum c_n \psi_{E_n}(x) e^{-\frac{i E_n}{\hbar} t}$$

* simulation *