

A white object and a black object both sit in an oven. The oven and the objects are in equilibrium at 1500 degrees Celcius. We can say that the **net** heat current from radiation,  $(H_{\text{absorbed}} - H_{\text{emitted}})$  is

- A) Larger for the white object
- B) Larger for the black object
- C) The same for both objects and greater than zero.
- D) The same for both objects and equal to zero.
- E) The same for both objects and less than zero.

*Assume that there are no conduction or convection effects.*

**EXTRA:** Which object is emitting more radiation?

# How to study for the midterm:

- ① Do practice problems
  - homework, tutorials, old midterms, clicker questions
  - use slides, solutions, problem solving tips as reference,
- ② Do more practice problems
- ③ Repeat

But continue to eat, sleep, exercise

Nobel Prize 2018 with Arthur Ashkin  
Gerard Mourou

eng-physics!

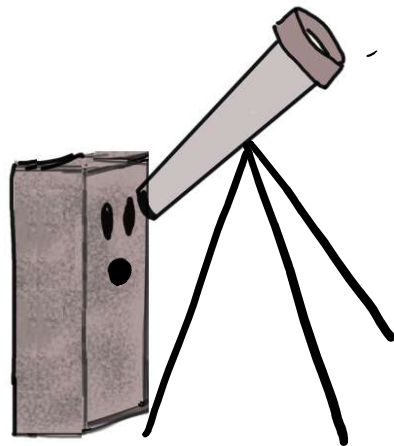
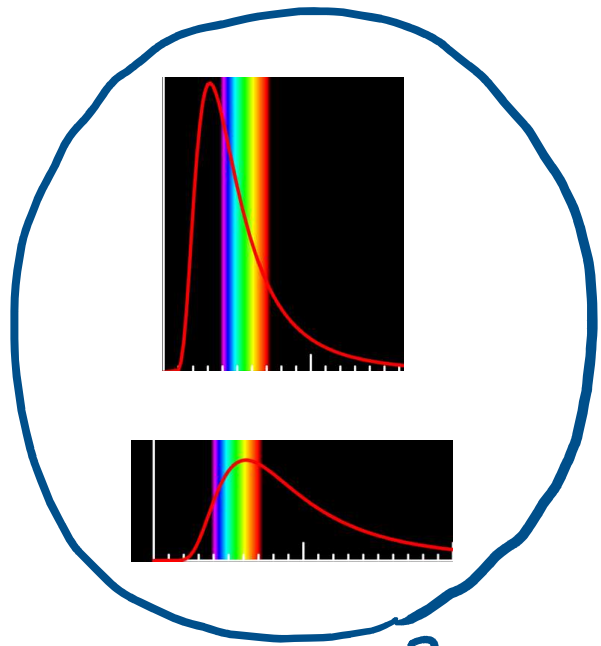
Donna T. Strickland - U. Waterloo



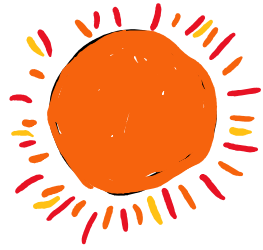
**2013 OSA President Donna Strickland** received her B. Eng. Degree in Physics, from McMaster University in 1981. She graduated from the University of Rochester in 1989 with a Ph.D. in Optics. Along with her Ph.D. supervisor, Dr. Gerard Mourou, Donna Strickland co-invented Chirped Pulse Amplification (CPA), which made it possible to amplify ultra-short pulses to unprecedented levels. From 1988 to 1991, Strickland was a research associate at the National Research Council of Canada. The following year, she was a physicist with the laser division of Lawrence Livermore National Laboratory. In 1992, she became a member of the technical staff of Princeton's Advanced Technology Center for Photonics and Opto-electronic Materials. Strickland joined the physics department of the University of

from  
Guelph,  
Ontario

Waterloo as an assistant professor in 1997. At Waterloo, Strickland's ultrafast laser group develops high-intensity laser systems for nonlinear optics investigations. She was promoted to Associate Professor in 2002 and since 2007 has been the Associate Chair of the Department. Strickland was selected as an Alfred P. Sloan Research Fellow in 1998. She received a Premier's Research Excellence Award in 1999 and a Cottrell Scholars Award from Research Corporation in 2000 and was named a Fellow of the OSA in 2008.



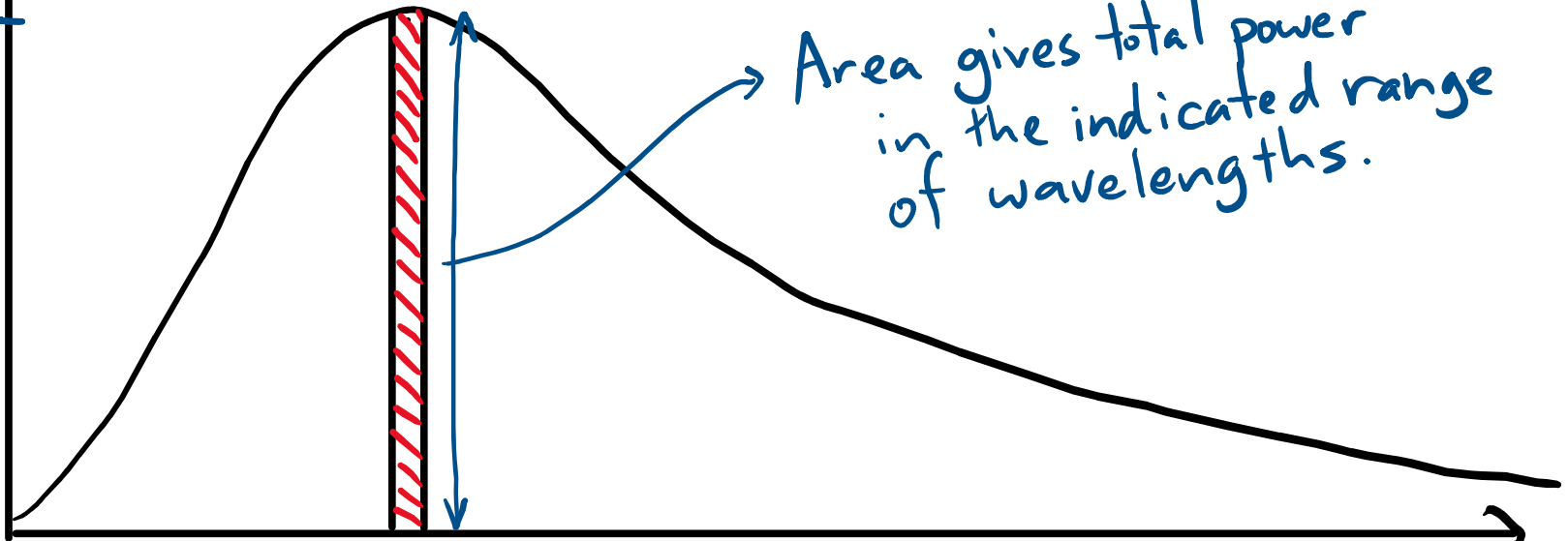
Last time  
in physics  
157...



Radiation from an object comes in a mix of wavelengths.  
We can describe this by graphing the SPECTRUM.

Power  
per nm

$400 \frac{\text{W}}{\text{nm}}$

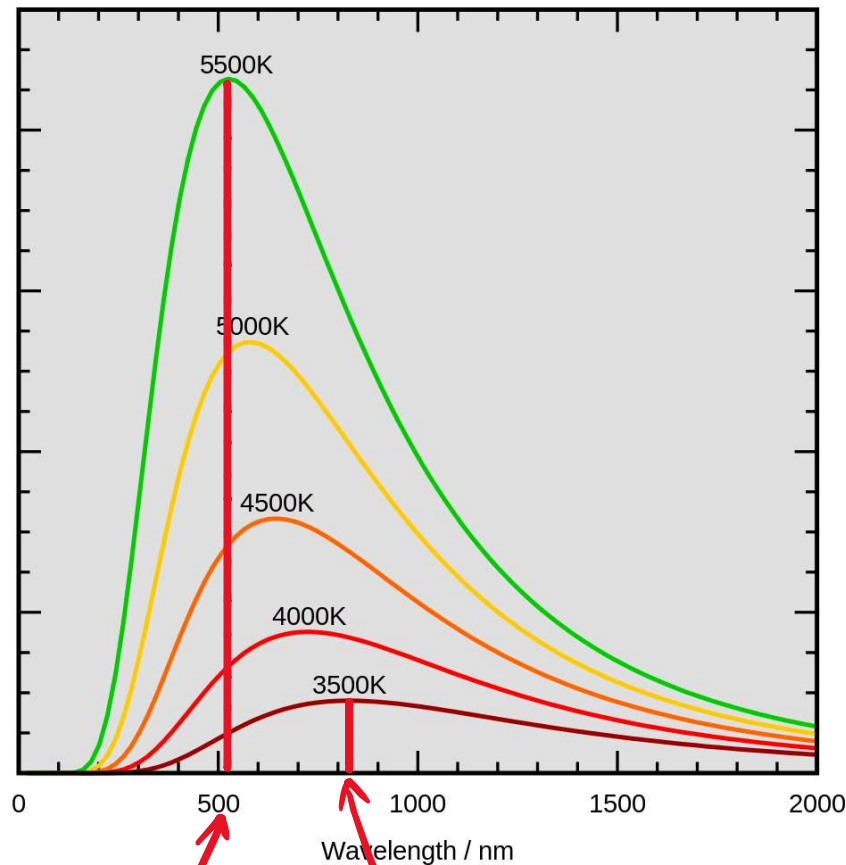


$20 \text{ nm}$

wavelength (nm)

Area gives total power  
in the indicated range  
of wavelengths.

Peak wavelength is inversely proportional to  $T$



$\lambda_{\max}$   
for  
5500K

$\lambda_{\max}$   
for  
3500K

$$\lambda_{\max} = \frac{b}{T}$$

2.9 K·mm

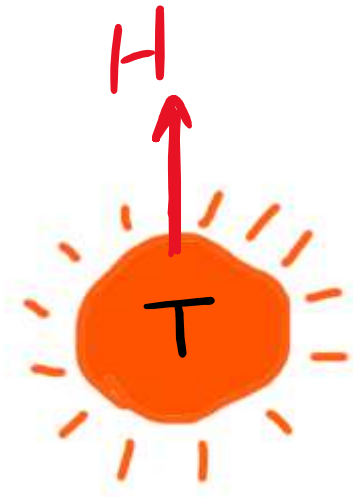
Wien displacement law

sun: peak at  $\approx 500\text{nm}$   
→ 5700K

outer space: peak at 1mm  
→ 2.7 K

"COSMIC MICROWAVE BACKGROUND"

Total power is proportional to  $T^4$



heat current

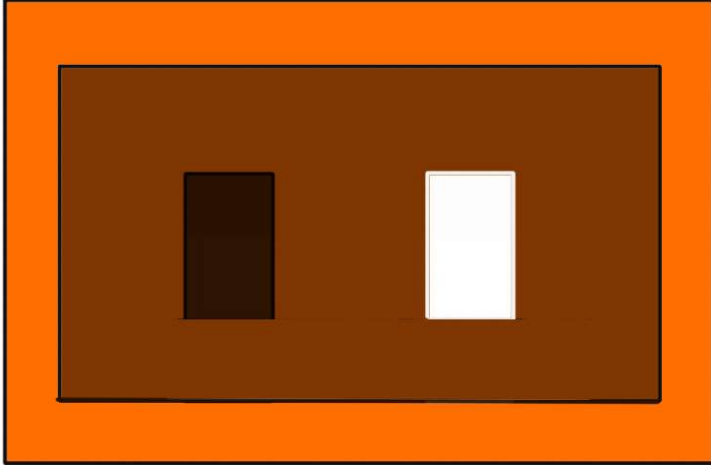
surface area

$$H = A \cdot e \cdot \sigma \cdot T^4$$

emissivity

Stefan-Boltzmann constant

$$5.67 \times 10^{-8} \frac{\text{W}}{\text{m}^2 \cdot \text{K}^4}$$



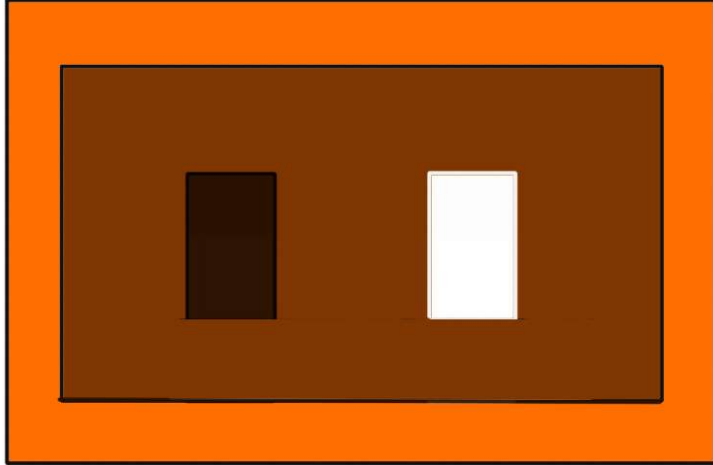
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*Assume that there are no conduction or convection effects.*

**EXTRA:** Which object is emitting more radiation?





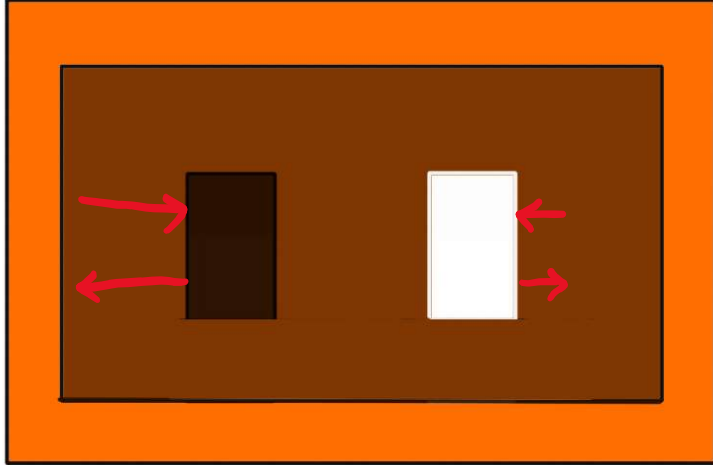
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Equilibrium  $\Rightarrow$  const.  $T$   
 $\Rightarrow$  no net heat current

$$\therefore H_{\text{absorbed}} - H_{\text{emitted}} = 0$$

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- B) Larger for the black object
- C) The same for both objects and greater than zero.
- D) The same for both objects and equal to zero.**
- E) The same for both objects and less than zero.

*Assume that there is no air in the oven and the objects are insulated from the walls so there is no conduction or convection.*



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$$H_{\text{emitted}} = H_{\text{absorbed}}$$



larger for black object

$\therefore$  black object radiates more!

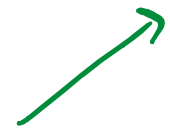
## EMISSIVITY:

- Perfect absorber = "blackbody" emits the most thermal radiation for a given temperature.

- Other objects: define

$$e = \frac{H}{H_{\text{blackbody}}}$$

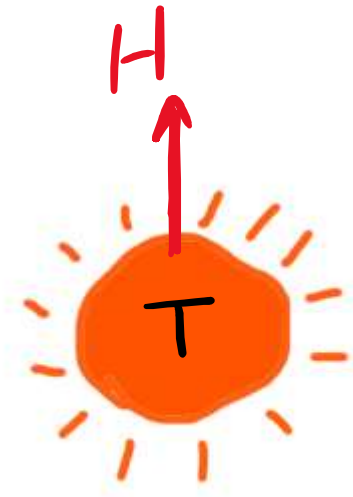
depends on temperature



$e = 1$  blackbody

$e = 0$  perfect mirror

# TOTAL POWER FROM THERMAL RADIATION



heat  
current

surface  
area

emissivity

$$H = A \cdot e \cdot \sigma \cdot T^4$$

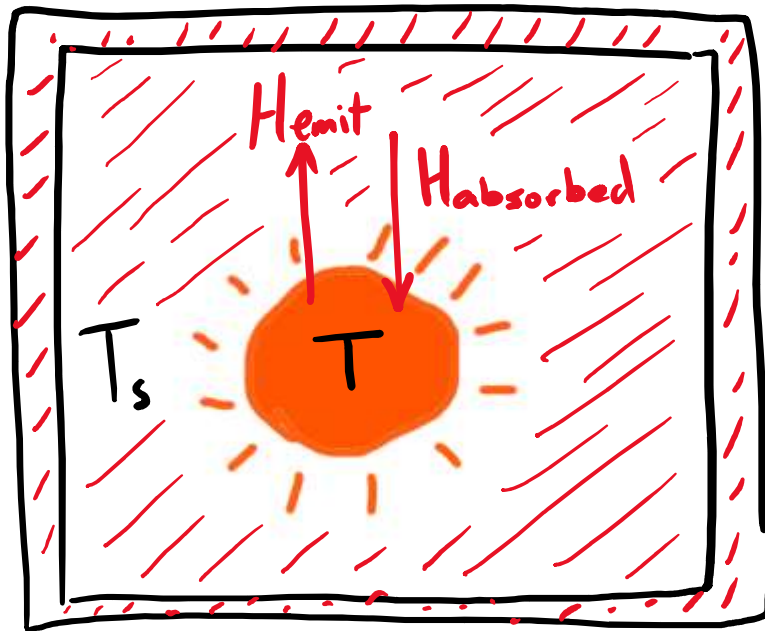
Stefan-  
Boltzmann  
constant

$$5.67 \times 10^{-8} \frac{\text{W}}{\text{m}^2 \cdot \text{K}^4}$$

$e = 1$  perfect  
absorber (black)

$e = 0$  perfect  
reflector (mirror)

# NET HEAT CURRENT FROM THERMAL RADIATION (in uniform temperature environment)



$e = 1$  perfect absorber (black)

$e = 0$  perfect reflector (mirror)

surface area

$$H = A \cdot e \cdot \sigma \cdot (T^4 - T_s^4)$$

Stefan-Boltzmann constant  
 $5.67 \times 10^{-8} \frac{W}{m^2 \cdot K^4}$

temp. of surroundings

Yoltar heats their little planet (far from any stars) with a 1GW heater. If they wish to double the equilibrium *surface* temperature of their planet, they should increase the power of their heater to

- A) 1.21GW
- B) 2GW
- C) 4GW
- D) 8GW
- E) 16GW



Hint: where does the energy from the heater go?



Steady state:

Power from heater  
= power radiated

$$P_{\text{heater}} = A \cdot \sigma \cdot e \cdot T^4$$

To double  $T$   
Need  $16 \times P$

A harder (but really interesting!) problem.

A planet with radius  $r = 6400\text{km}$  lies at a distance  $R = 150,000,000\text{km}$  from a yellow star with temperature  $T = 5700\text{K}$  and radius  $R_s = 695,000\text{km}$ . **Estimate the surface temperature of the planet.**

The planet has **albedo** (fraction of incident light reflected)  $A = 0.37$  and emissivity  $e$  close to 1.



STEP 1: visualize what is going on in the problem.  
- draw a picture

Discussion question: what physical processes are happening here that will affect the the temperature of the planet?

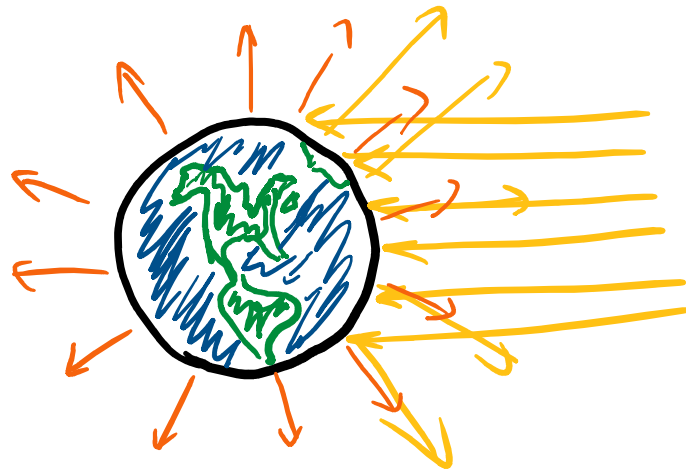
If the temperature is constant, what does this allow us to conclude?

Key relation for steady-state heat flow:

$$H_{in} = H_{out}$$



Our problem:



$H_{in}$ : absorbed  
sunlight

$H_{out}$ : IR radiation  
 $= A \cdot e \cdot \sigma \cdot T^4$

What is  $H_{in}$ ?