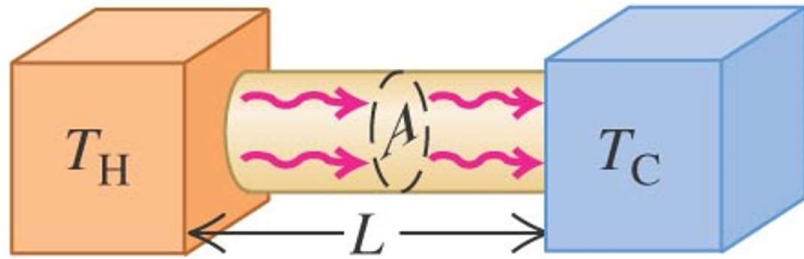


THERMAL CONDUCTIVITY: Determines heat current from temperature gradient.



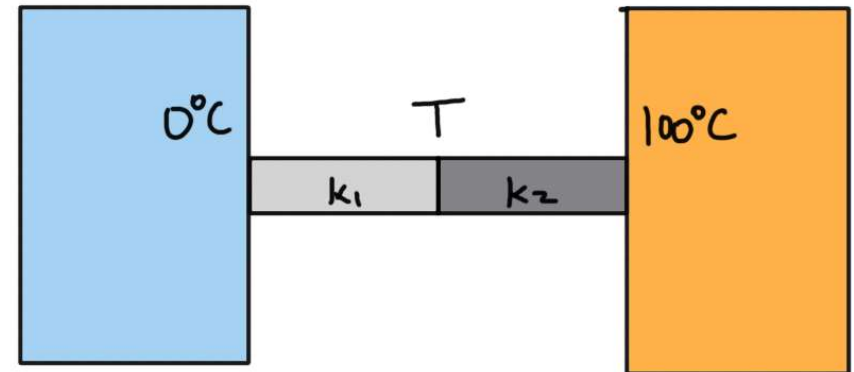
$$H = k A \frac{T_H - T_C}{L} \left\{ \begin{array}{l} \text{temperature} \\ \text{gradient} \end{array} \right.$$

Heat current
"
Heat per time

Thermal
conductivity

Homework sessions: Monday 5-7pm, Hennings 200
Tuesday 5-7pm, Hennings 202

Two materials of equal dimensions but different thermal conductivities are placed side to side between objects kept at 0°C and 100°C , and a steady heat flow is established. If $k_1 < k_2$, we can say that the temperature T in the middle is:



- A) Equal to 50°C B) Greater than 50°C C) Less than 50°C

EXTRA: How would you calculate the temperature.

Homework sessions: Monday 5-7pm, Hennings 200
Tuesday 5-7pm, Hennings 202

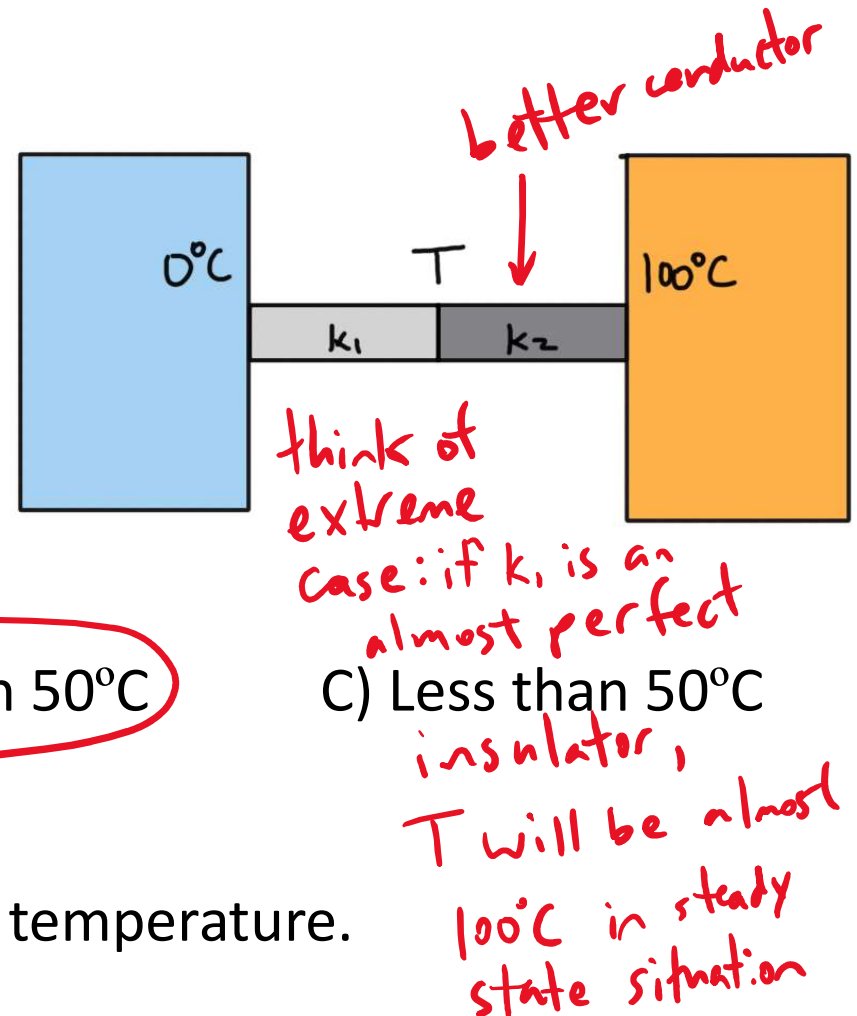
Two materials of equal dimensions but different thermal conductivities are placed side to side between objects kept at 0°C and 100°C , and a steady heat flow is established. If $k_1 < k_2$, we can say that the temperature T in the middle is:

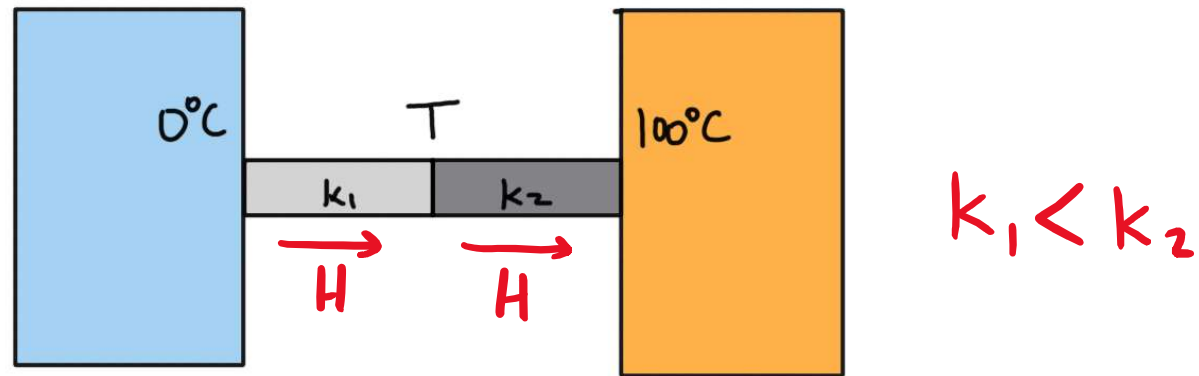
A) Equal to 50°C

B) Greater than 50°C

C) Less than 50°C

EXTRA: How would you calculate the temperature.

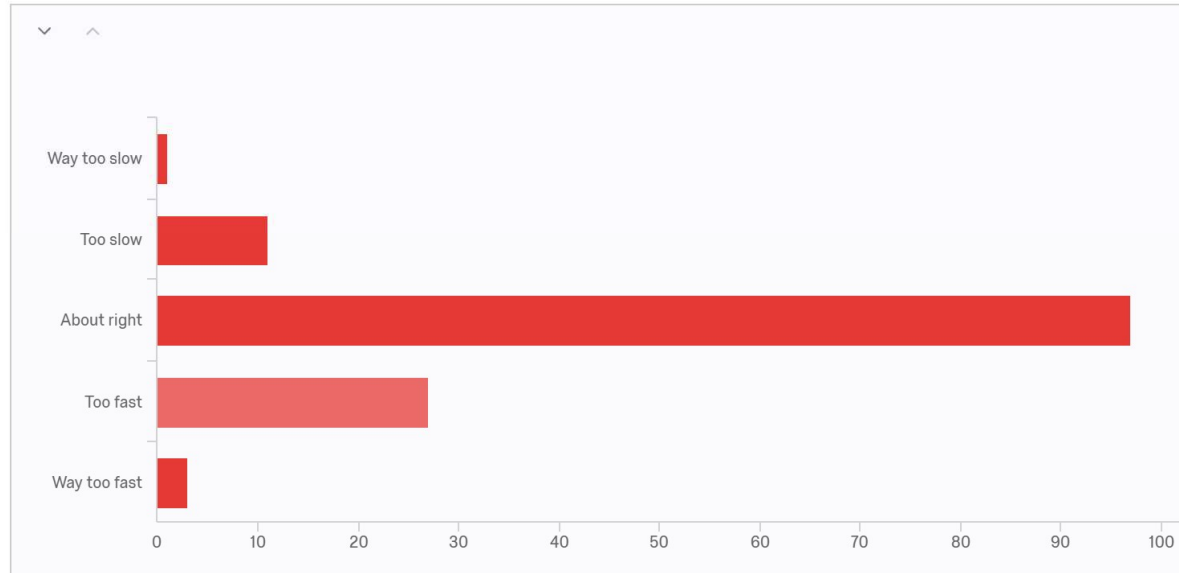




- H is the same for both parts.
- $H = k_1 \cdot A \cdot \frac{T - 0^\circ\text{C}}{L} = k_2 \cdot A \cdot \frac{100^\circ - T}{L}$
- $k_1 < k_2$ so $T - 0^\circ\text{C} > 100^\circ\text{C} - T$: T is closer to 100°

Q1 - The pace of the course is:

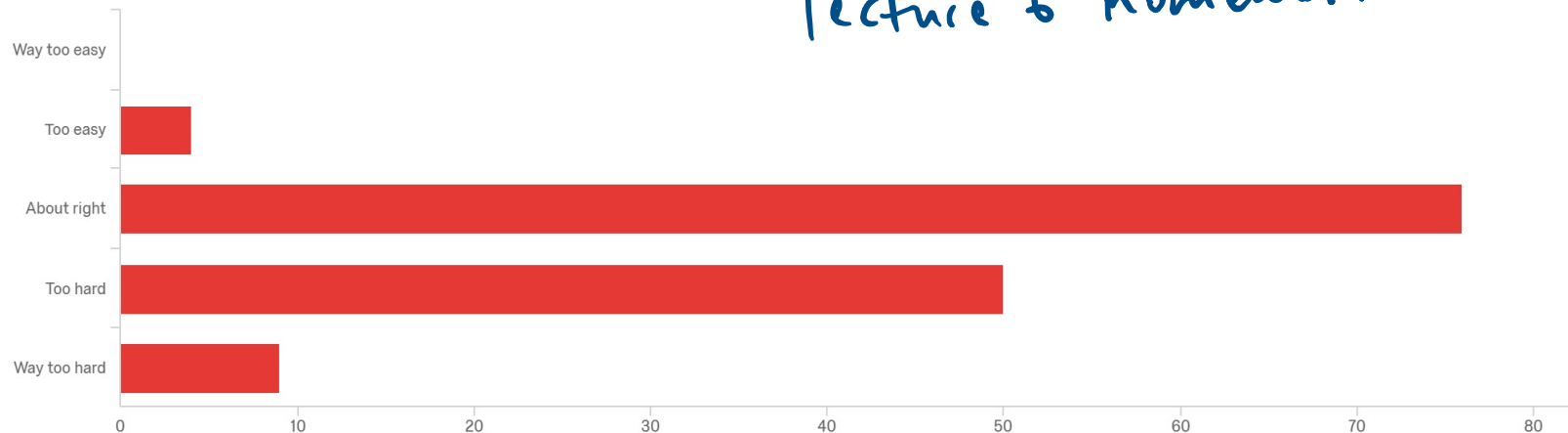
Page Optic



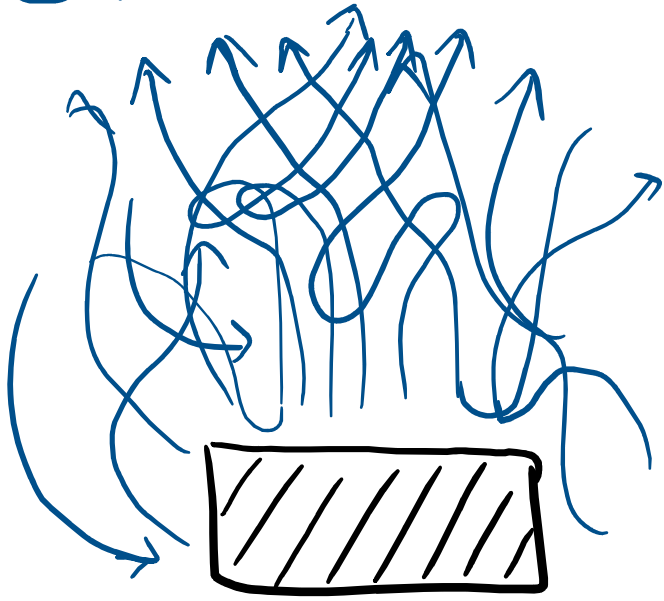
Q2 - The homework is:

top Comments: tutorials could be improved
- more direct connection between
lecture & homework.

Page Optic

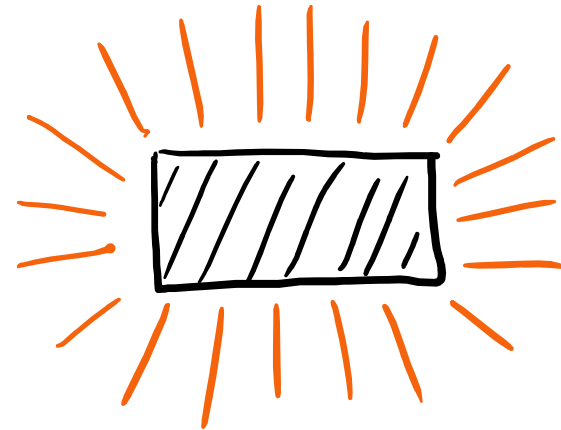


Other mechanisms for heat transfer:



CONVECTION: heat transfer via macroscopic motion of fluids

- very complicated fluid dynamics to understand

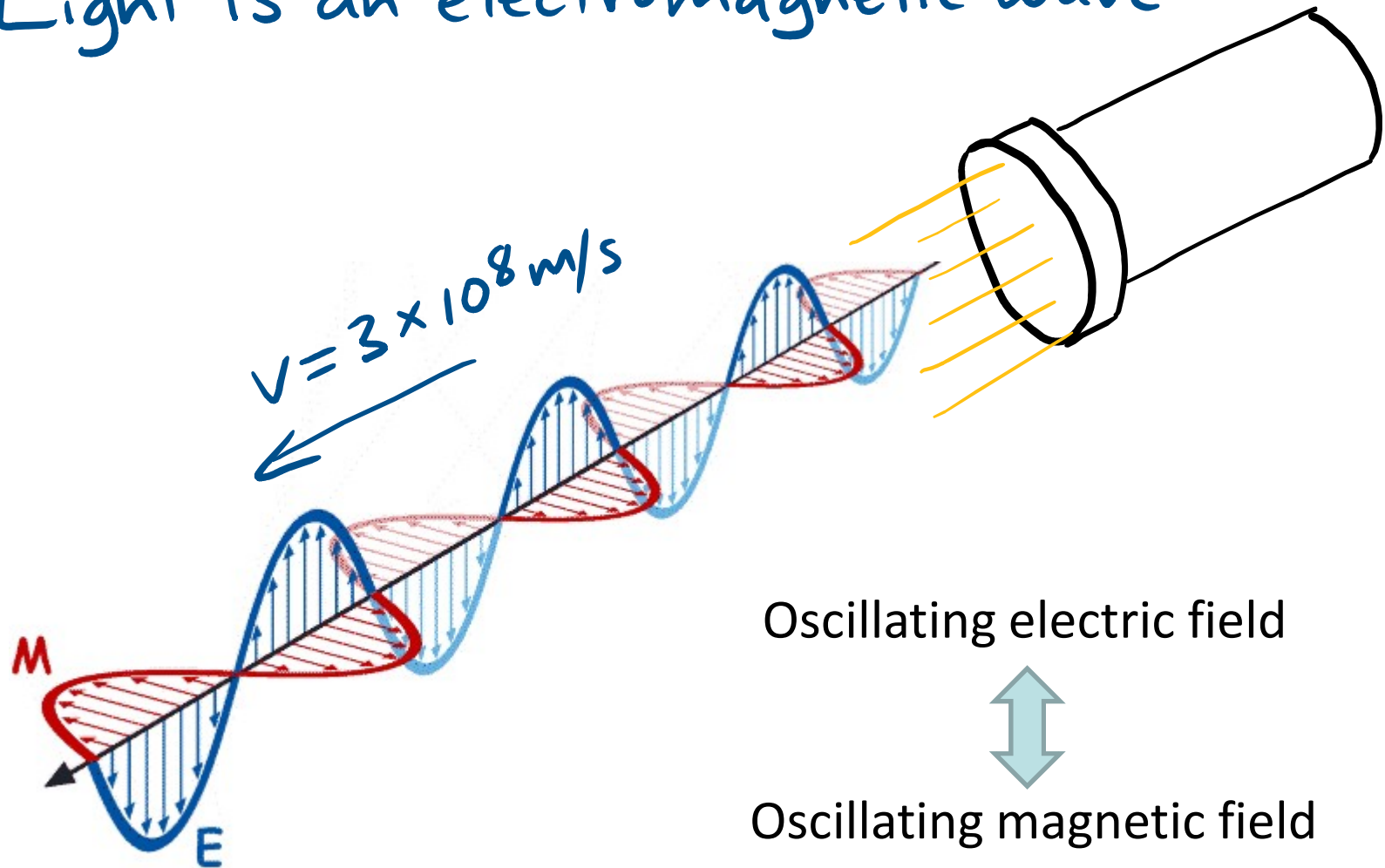


RADIATION: all objects give off electromagnetic radiation (light, IR, etc...)

- this carries energy away from the object

ELECTROMAGNETIC RADIATION:

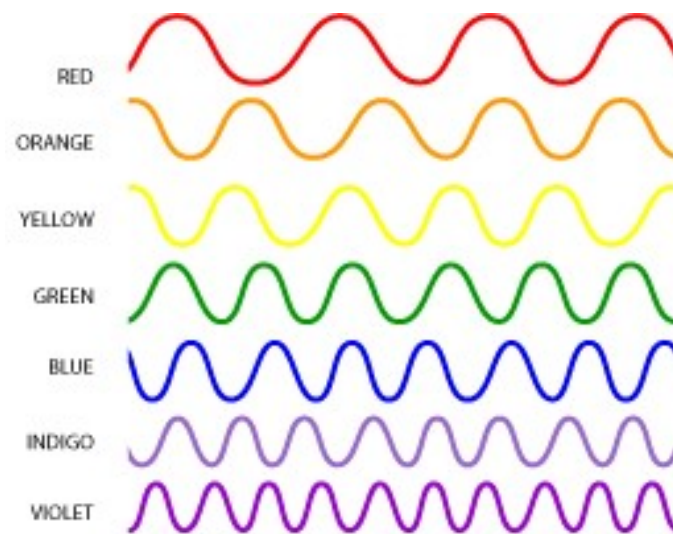
Light is an electromagnetic wave:



James Clerk Maxwell 1864

Properties of Light

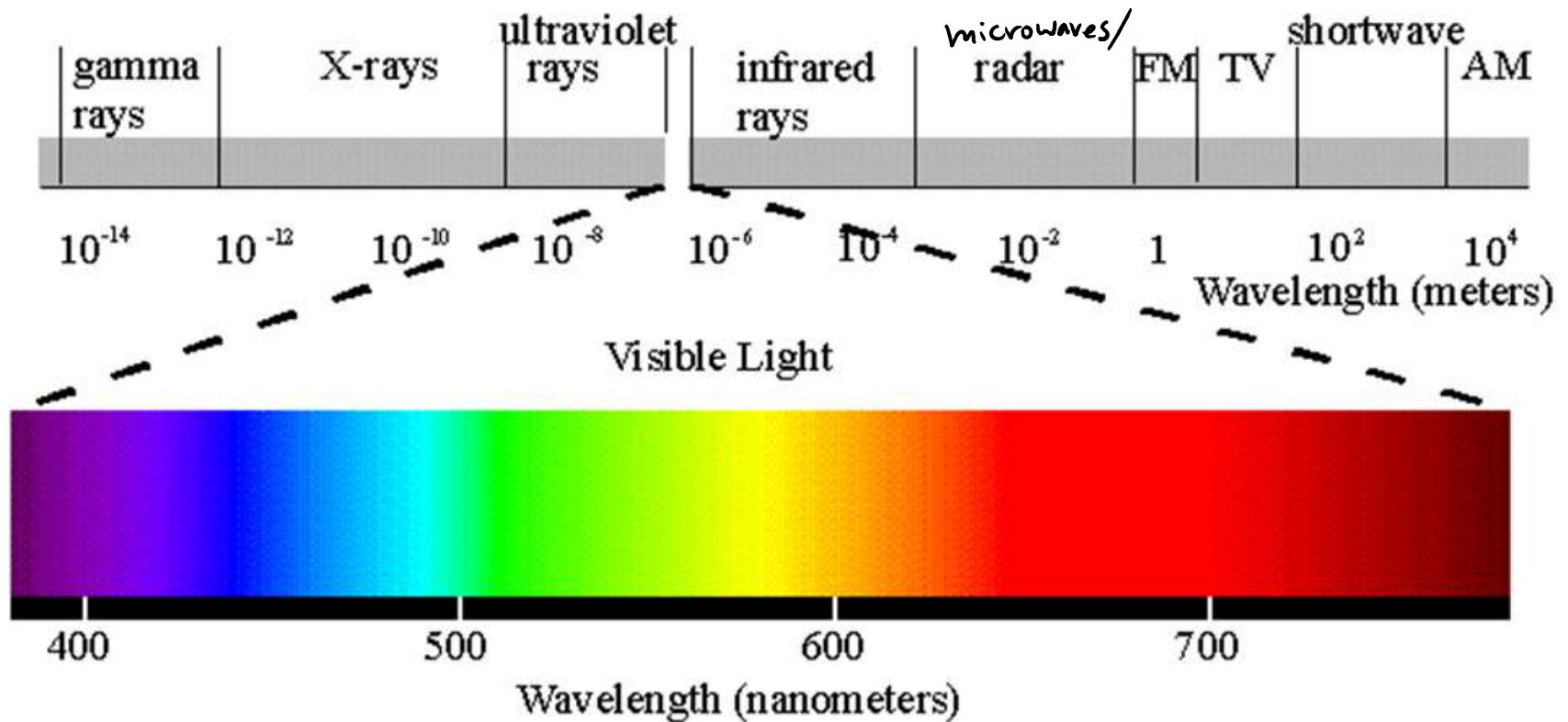
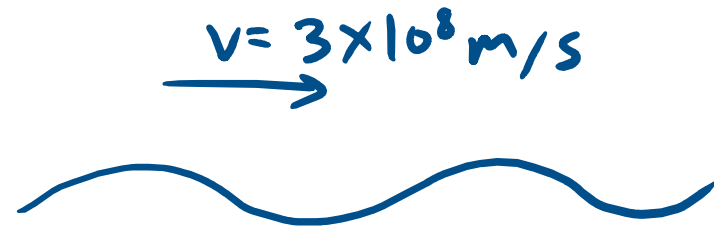
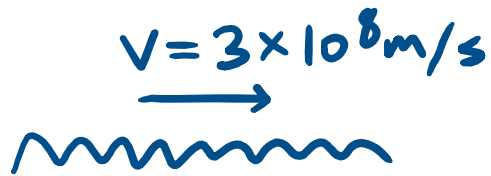
Colour: determined by wavelength



Intensity/brightness: determined by amplitude

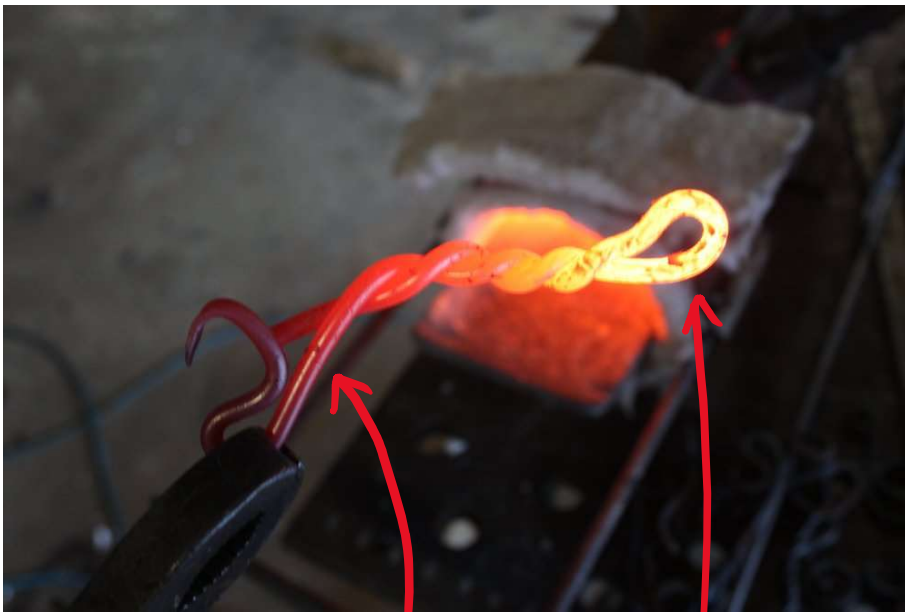
↳ light carries energy!

Can have electromagnetic waves at all wavelengths:



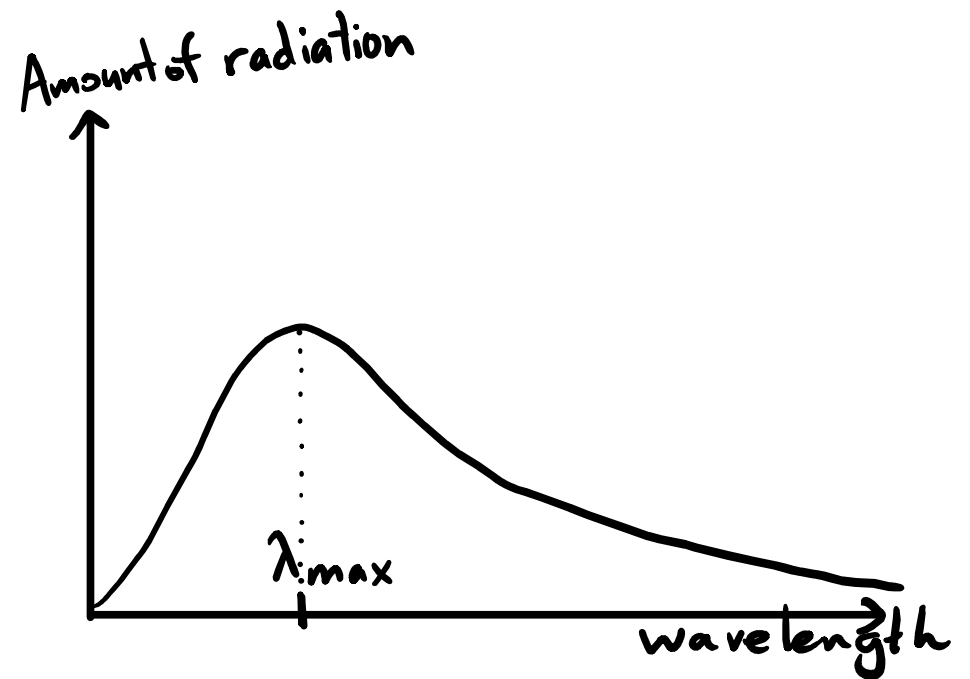
Thermal radiation from an object:

- typically in IR/visible
- can measure energy current at various wavelengths = spectrum

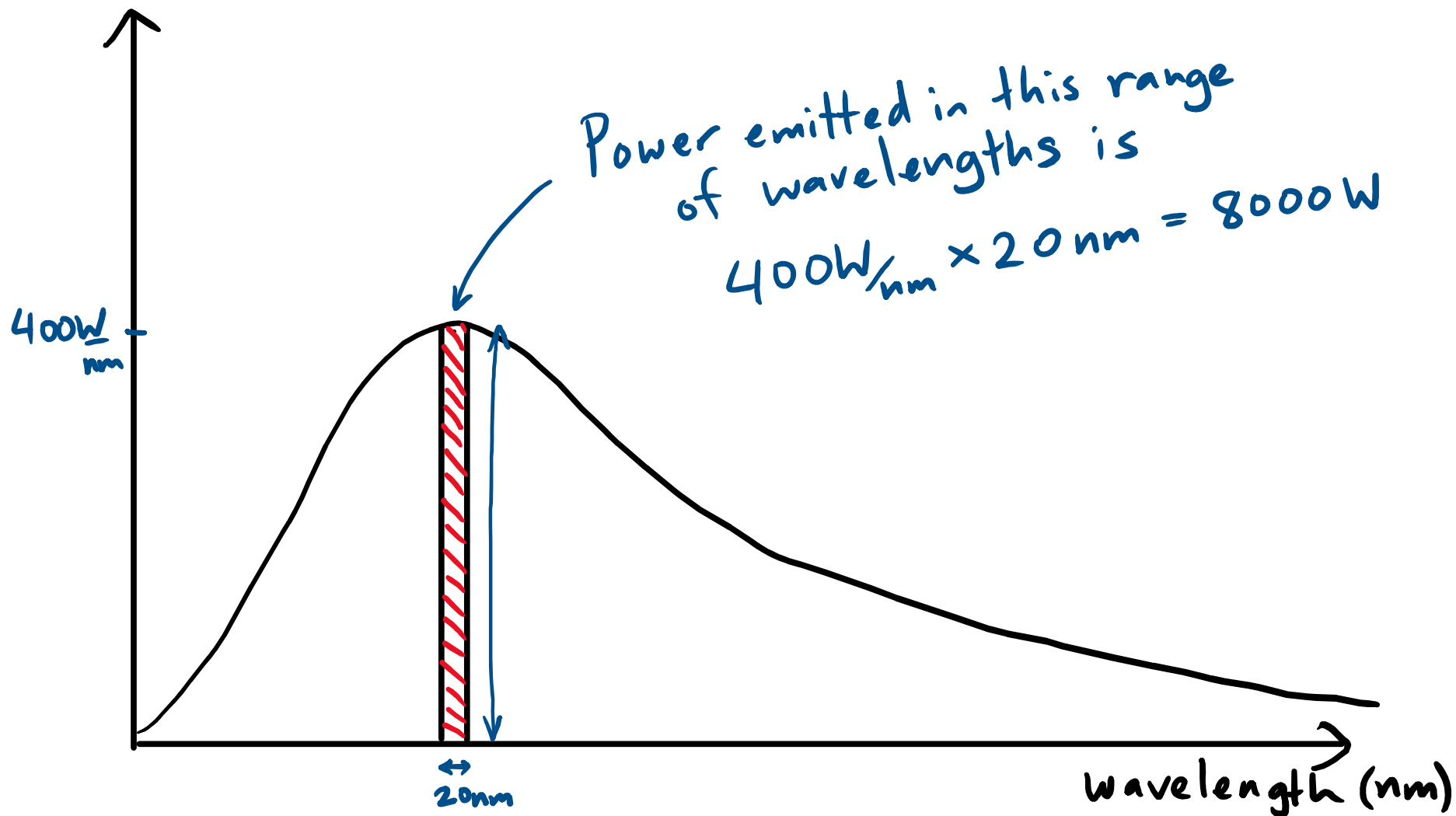


cooler

hotter

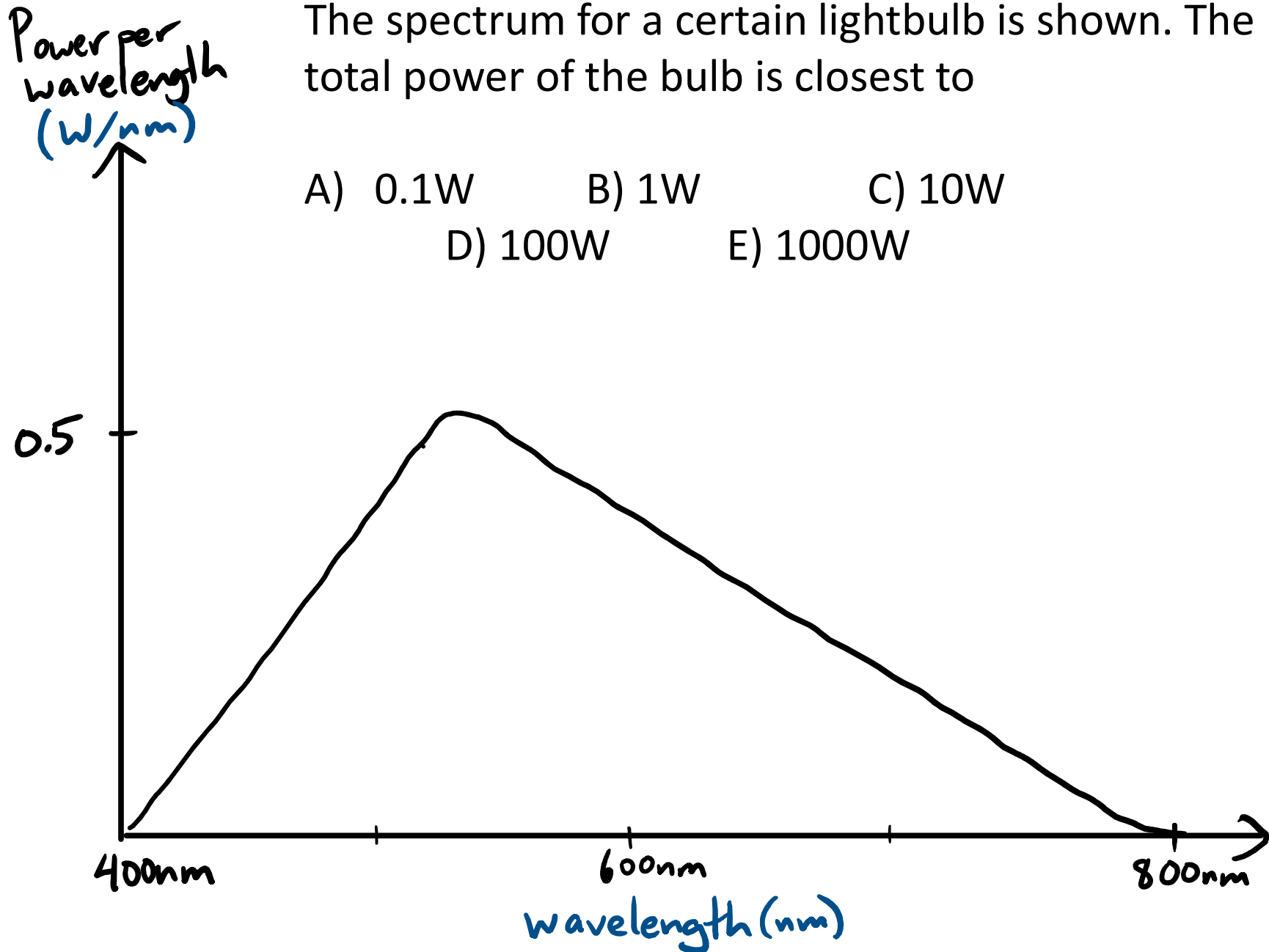


Power
per nm



The spectrum for a certain lightbulb is shown. The total power of the bulb is closest to

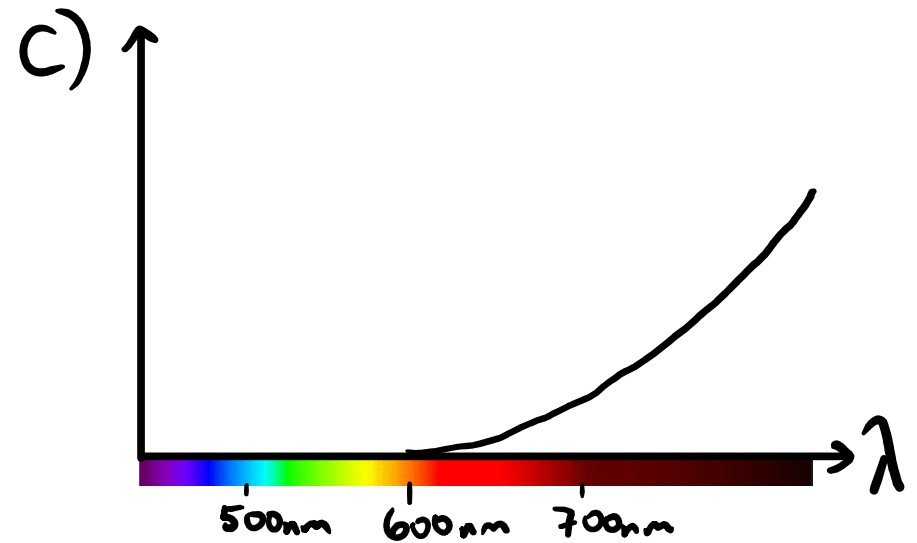
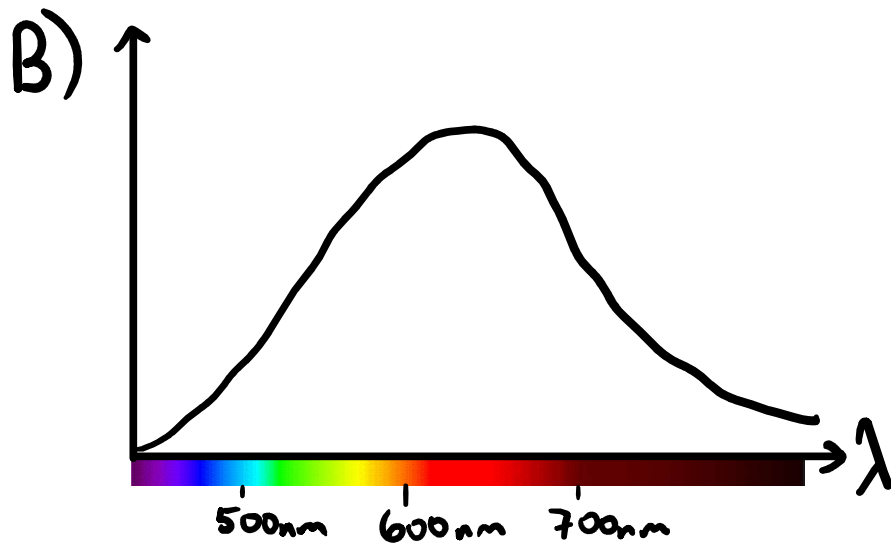
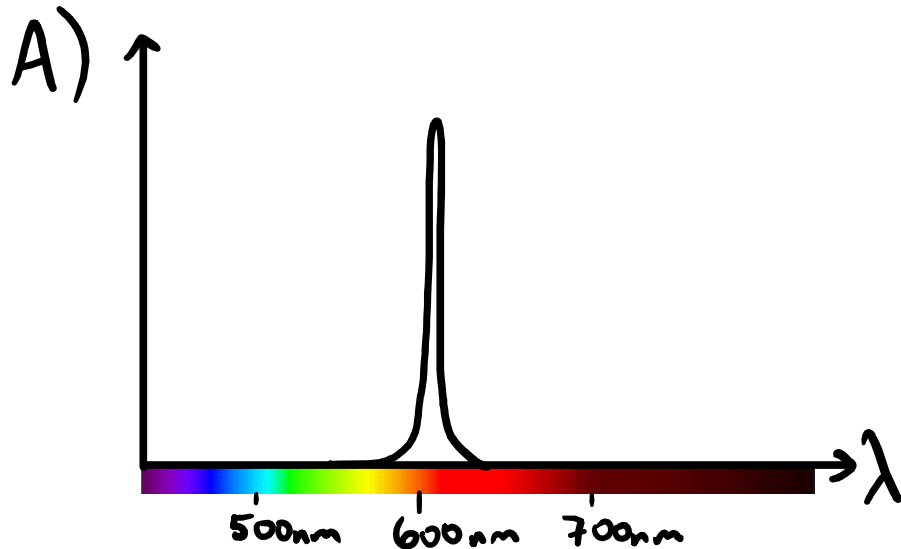
- A) 0.1W B) 1W C) 10W
D) 100W E) 1000W



Which graph best represents the spectrum of radiation from the red hot ball in the picture?



from Youtube: 1000 Degree Metal Ball vs Milk

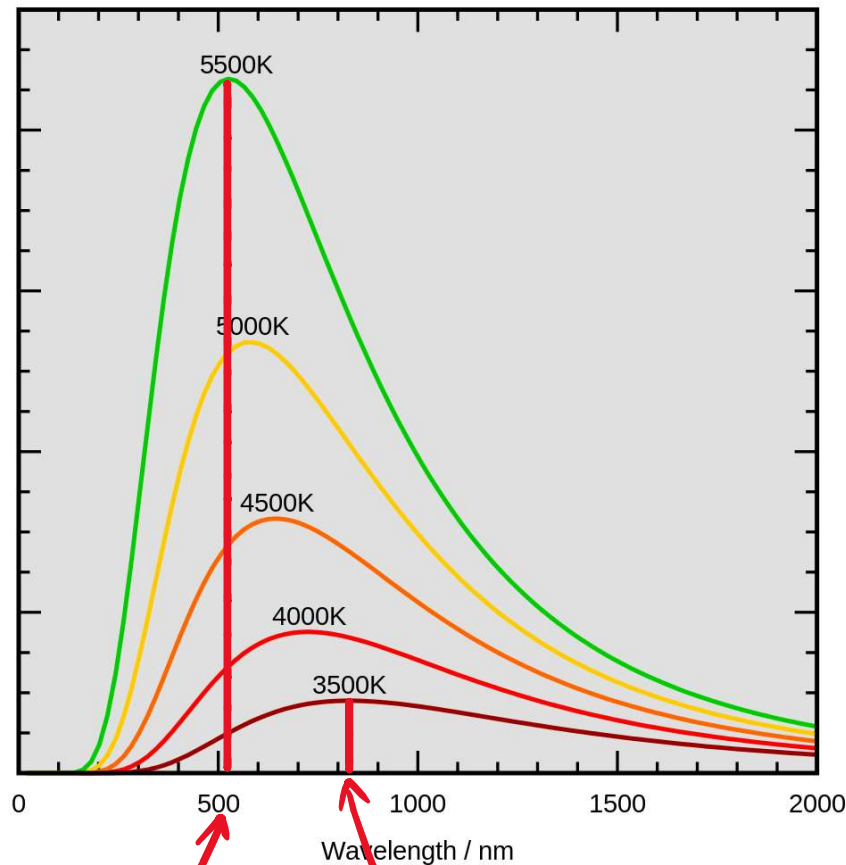


Blackbody spectrum

https://phet.colorado.edu/sims/blackbody-spectrum/blackbody-spectrum_en.html

In the simulation, what properties of the thermal spectrum change as we change the temperature?

Peak wavelength is inversely proportional to T



λ_{\max}
for
5500K

λ_{\max}
for
3500K

$$\lambda_{\max} = \frac{b}{T}$$

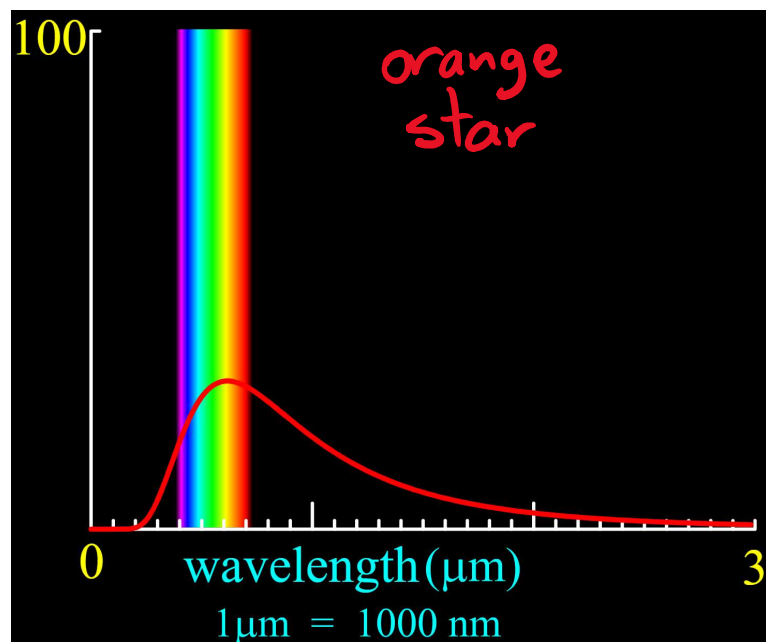
2.9 K·mm

Wien displacement law

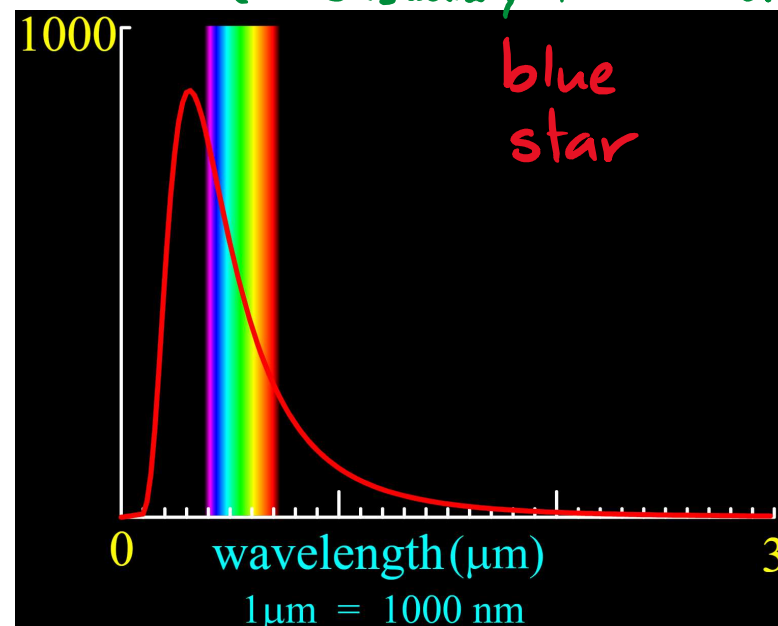
sun: peak at $\approx 500\text{nm}$



(there is actually a third star we're ignoring!)

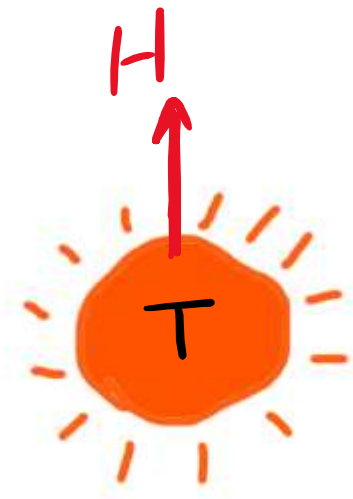


$T \approx 4500\text{K}$



$T \approx 12,000\text{K}$

TOTAL POWER FROM THERMAL RADIATION



heat
current

surface
area

emissivity

$$H = A \cdot e \cdot \sigma \cdot T^4$$

Stefan-
Boltzmann
constant

$$5.67 \times 10^{-8} \frac{\text{W}}{\text{m}^2 \cdot \text{K}^4}$$

$e = 1$ perfect
absorber (black)

$e = 0$ perfect
reflector (mirror)