

An informal, community-organized conference borne from the desire for people to share and learn in an open environment. Everyone from casual skeptics to the experienced participate, give talks, and get to know each other.

Completely Free

vancouver.skepticamp.org

10am-6pm Saturday, March 20, 2010 Victoria Learning Theatre (Room 182) Irving K. Barber Learning Centre 1961 East Mall, UBC, Vancouver

http://vancouver.skepticamp.org/

Happy belated birthday, Albert!

Albert Einstein was born on 14 March 1879 in Ulm, Germany









Be careful if you Google this phrase







The user car lot





Do I have a <u>deal</u> for you!

The user star lot In some ways, stars are like cars



The user star lot In some ways, stars are like cars



<u>Massive stars</u> = <u>sports cars</u>





Hot massive stars are bright and attract attention across great distances

<u>Massive stars</u> = <u>sports cars</u>





Hot massive stars are bright and attract attention across great distances but they're not very fuel efficient

<u>Solar-mass stars</u> = <u>economy cars</u>





2009 model Sol

Stars of moderate mass are moderately bright and much more fuel efficient

<u>M dwarf stars</u> = <u>hybrid super-economy cars</u>





The lowest-mass stars are hard to notice unless they're very close by

<u>M dwarf stars</u> = <u>hybrid super-economy cars</u>







The lowest-mass stars are hard to notice unless they're very close by

but boy, do they go a *long way* on a single 'tank' of hydrogen

The showroom floor



B







G

F

Μ

K

Main sequence energy generation



Main sequence energy generation



Main sequence lifetime



mass-luminosity relation $L/L_{Sun} \sim (M/M_{Sun})^{3.5}$

Main sequence lifetime



mass-luminosity relation $L/L_{Sun} \sim (M/M_{Sun})^{3.5}$ total energy = luminosity × time $t/t_{Sun} \sim (E/E_{Sun})/(L/L_{Sun})$

<u>Main sequence lifetime</u>



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Main sequence lifetime



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Main sequence lifetime

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How does a star produce energy?

Through thermonuclear fusion of H to He for most of its life How much energy does this produce?



where ΔM is the total amount of mass destroyed

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The total amount of hydrogen fuel available to a star is proportional to its total mass

 $\Delta M = bM$

where b is the fraction of mass consumed

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B2 dwarf

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> If a star is 1 tenth the mass of the Sun it will be able to shine for (0.1)^{-2.5} ~ <u>300 times as long</u>

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The total lifetime of the Sun is estimated to be about 10 billion (10¹⁰) years

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M8 dwarf

B2 dwarf



An M dwarf lasts a *long* time on a single tank of hydrogen









B



A stable, "boring" life on the main sequence

While on the main sequence stars are in hydrostatic equilibrium so changes are slow and subtle



A stable, "boring" life on the main sequence

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> As hydrogen in the stellar core is gradually being replaced by helium the star's overall equilibrium also changes

For example, the Sun has been very slowly increasing in radius and in luminosity during its main sequence life

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> As hydrogen in the stellar core is gradually being replaced by helium the star's overall equilibrium also changes

For example, the Sun has been very slowly increasing in radius and in luminosity during its main sequence life

In 4½ billion years since the Sun was born it has increased in luminosity by a few tens of per cent and it will continue to brighten for the remainder of its life

The Faint Young Sun Paradox

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The Faint Young Sun Paradox

Astronomers know that the Sun has been gradually increasing in luminosity over the history of the Solar System so that the Earth must have been cooler in the past

But evolutionary models of the Sun predict that the solar luminosity was low enough in the past that the Earth should have been frozen over at the time when we know liquid oceans first appeared on our planet's surface

The Faint Young Sun Paradox



The Faint Young Sun Paradox

It is unlikely that the models of solar evolution are wrong by enough to resolve the paradox

The details of the Earth's primordial atmosphere and the early Greenhouse Effect are more promising ways to explain the Earth's relatively stable temperature history





<u>The past – Earth's first oceans</u>

Results announced today show that analysis of sandstones in North Greenland confirm that oceans formed on Earth about 4 billion years ago

http://www.cbc.ca/ technology/story/ 2010/03/15/ tech-geologyancient-ocean.html



<u>The future</u>

Even before the Sun exhausts its core hydrogen fuel supply it will continue to expand and to brighten slowly



<u>The future</u>

Even before the Sun exhausts its core hydrogen fuel supply it will continue to expand and to brighten slowly





This will cause the Earth's surface temperature to rise In a few 100 million years to a billion years or so the oceans will boil away

Questions?

Travelling on the MOST transit system









Transiting exoplanet



artist's conception

Sun

Mercury

Transiting exoplanet



Position: $RA = 22\ 03\ 10.8$ Dec = +18 53 04 Distance: 47 pc ~ 153 ly Constellation: Pegasus

Star

Magnitude (Star): V=7.64 Spectra Class: G0 V Temperature: 6000 K

Planet

Orbital Period: 3.52 days Semi-major axis: 0.046 AU Eccentricity: 0 Mass: 0.68 × Jupiter Radius: 1.35 × Jupiter



measurement of the transit of the giant planet by MOST



✓ measurement of the <u>eclipse</u> of the giant planet by MOST



Imagine trying to see a mosquito disappearing behind a 400-Watt streetlamp.



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Not at the street corner, nor a few blocks away...



Imagine trying to see a mosquito disappearing behind a 400-Watt streetlamp.

Not at the street corner, nor a few blocks away... but <u>1000 km</u> away.



Photon scattering & molecular absorption are the dominant mechanisms that determine the reflected spectrum of an exoplanet

MOST instrument throughput

 ✓ For wavelengths *λ* < 600 nm Rayleigh scattering dominates

 a large fraction of incident light can be reflected outwards

✓ For wavelengths *λ* > 600 nm photons are absorbed deep in the exoplanet's atmosphere



 Strong UV flux produces haze darkening planet in the blue
Cloud formation is still speculative

Particle size and composition?



Albedo, Bond Albedo



Bond Albedo



Bond Albedo

Phase function



Treating the phase function of the planet as an isotropic scatterer gives an upper limit on the Bond Albedo

MOST optical







Spitzer infrared Deming et al. 2005 Nature 111, 111

Rowe et al. 2008 Astrophysical Journal

MOST optical



Rowe et al. 2008 Astrophysical Journal





Spitzer *infrared* Deming et al. 2005 Nature 111, 111

Best fit parameters:

 $albedo = 0.04 \pm 0.04$

stellar radius : $1.339 \pm 0.001 R_{Jupiter}$ stellar mass $1.084 \pm 0.005 M_{Sun}$

> i = 86.937° ± 0.003° P = 3.5247489 d



Geometric Albedo













HD 209458 b from Earth



HD 209458 b from Earth



Exoplanetometeorology



MOST data are helping us understand the weather and clouds on a planet you can't even see around a star 160 light years away!?!