

Phys529B: Topics of Quantum Theory

Lecture 14: identifying Non-Fermi liquids via RGE

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Supplementary stuff on RGE

$$H = H(m, \lambda, \dots; \Lambda) = H(\tilde{m}, \tilde{\lambda}, \dots, Z(\Lambda); \Lambda)$$

for any given Λ , there shall be a $\tilde{m}(\Lambda), \tilde{\lambda}(\Lambda) \dots$

so that H leads to the same physics.

$$\frac{d\tilde{m}}{dt} = \beta_m(\tilde{m}, \tilde{\lambda}, \dots)$$

$$\frac{dZ(\Lambda)}{d\Lambda} = \gamma(\tilde{m}, \tilde{\lambda}, Z)$$

$$\frac{d\tilde{\lambda}}{dt} = \beta_{\tilde{\lambda}}(\tilde{m}, \tilde{\lambda}, \dots)$$

$$t = \ln \frac{\Lambda}{\Lambda_{uv}}$$

Renormalization Group Equations

Running scale

Generally $G = G_0 + G_0 \Sigma G$ or $G^{-1} = G_0^{-1} - \Sigma$

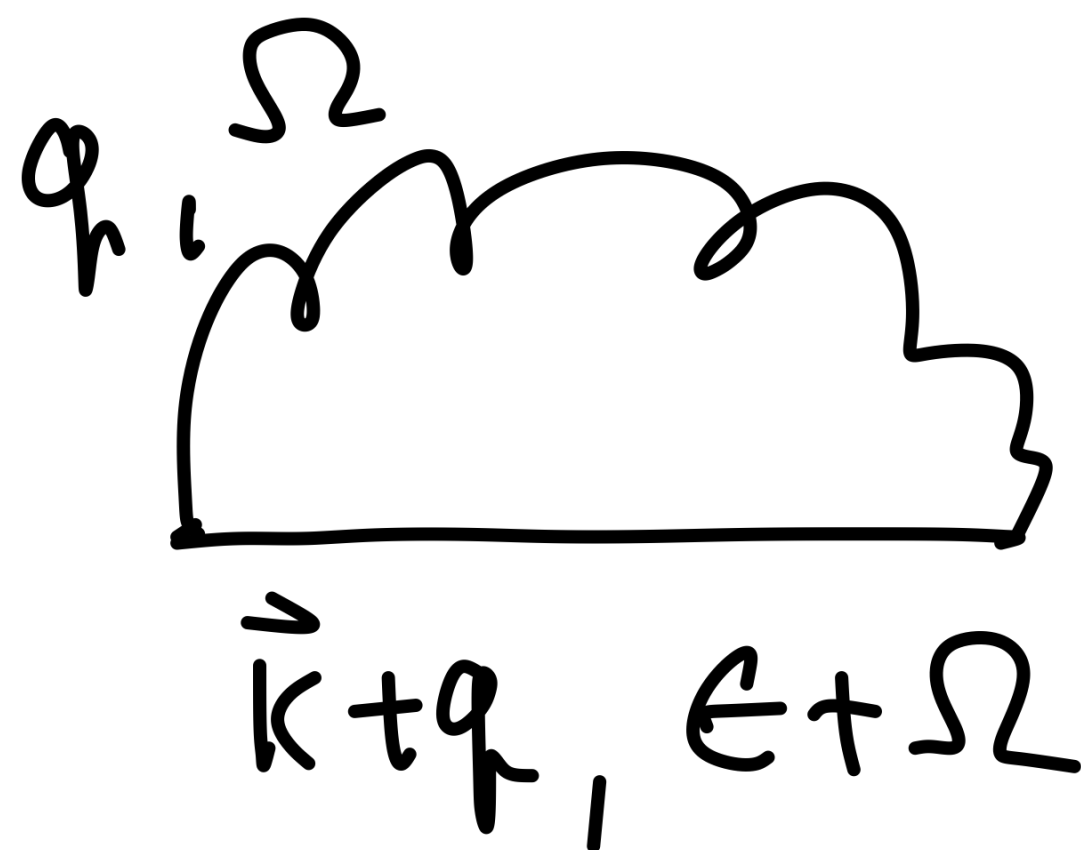
perturbation theory:



$$G_k = \langle 0 | -i T \psi_k(0) \psi_k^\dagger(0) | 0 \rangle$$

$$\delta G = G_0 \Sigma^{(2)} G_0,$$

$$\Sigma^{(2)}(\vec{k}, \epsilon) = i \int D_0(\vec{q}, \Omega) G_0(\vec{k} + \vec{q}, \epsilon + \Omega) \times g^2$$



$$D_0(\vec{q}, \Omega) = \frac{1}{\Omega^2 - \omega_{\vec{q}}^2 + i\delta}$$

$$G_0(\vec{k}, \epsilon) = \frac{1}{\epsilon - (\epsilon_{\vec{k}} - \mu) + i\delta_{\vec{k}}}$$

- The basic idea of renormalizability: 1) physical measurable shall be independent of the UV scale at which you formulate your theory; i.e. “scale independent”. 2) there can be infinite numbers of field theories of one single physical reality; they are all equivalent. 3) different theories is connected by Renormalization group equations.

What to do beyond perturbation theory: RGE
 idea

Only Λ , Not Λ_{uv}

$$P = P(g; \Lambda) = \Lambda^P \tilde{P}(\tilde{g}), \quad H = H(g; \Lambda)$$

$$0 = \frac{\partial P}{\partial t} = P P + \Lambda^P \frac{\partial \tilde{P}(\tilde{g})}{\partial \tilde{g}} \beta(\tilde{g}), \quad \beta(\tilde{g}) = \frac{\partial \tilde{g}}{\partial t}$$

$$\beta(\tilde{g}) = -P \frac{\tilde{P}}{\frac{\partial P}{\partial \tilde{g}}} \leftarrow \text{a function of } \tilde{g}.$$

Supplementary stuff on RGE

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Renormalization Group Equations

Running scale

$$\Lambda_{uv} \quad \phi_0(x) \quad H(g_{01}, g_{02}, \dots; m, \Lambda_{uv})$$

$$\Lambda \quad \phi(x) \quad H(g_1(\Lambda), g_2(\Lambda), \dots; m(\Lambda), \Lambda; z(\Lambda))$$

$$\phi(x) = \sum^{-\frac{1}{2}} \phi_R(x)$$

$$\text{or} = \sum^{\frac{1}{2}} \phi_0(x)$$

$$\Lambda_{2R} \quad \phi_R(x) \quad H(g_{1R}, g_{2R}, \dots; m_R)$$

- Callan-Symanzik approach (also Coleman-Weinberg applications) : Utilizing Green's functions to obtain RGEs and understand scale Transformations
- An interesting application to cold gases can be found in Yang, Jiang and Zhou, PRL 124 (22), 225701, 2020, Tricritical physics in 2D Superfluids.

Approach 1: $\phi(x) = Z^{\frac{1}{2}}(\Lambda) \phi_R(x)$

$$\langle 0 | T \phi(x_1) \phi(x_2) \dots \phi(x_n) | 0 \rangle = G^{(n)}(x_1, \dots, x_n)$$

$$G^{(n)}(x_1, \dots, x_n) = Z^{n/2}(\Lambda) \underbrace{\langle 0 | T \phi_R(x_1) \dots \phi_R(x_n) | 0 \rangle}_c$$

independent of " Λ ".

Computed with $H(g_1(\Lambda), g_2(\Lambda), \dots, g_n(\Lambda); m(\Lambda), \Lambda)$

$$G^{(n)} = G^{(n)}(\dots; \hat{g}_1, \hat{g}_2, \dots, \hat{g}_n(\Lambda); \hat{m}(\Lambda), t)$$

$$t = \ln \Lambda$$

$$\left\{ \frac{\partial}{\partial t} + \beta(\vec{g}) \frac{\partial}{\partial \vec{g}} \right\} G^{(n)} = \left\{ \frac{n}{2} \underbrace{\frac{\delta Z}{\delta \Lambda}}_{\gamma(\vec{g})} \frac{1}{Z} \right\} G^{(n)}$$

$$\left\{ \frac{\partial}{\partial t} + \beta(\vec{g}) \frac{\partial}{\partial \vec{g}} - \gamma(\vec{g}) \right\} G^{(n)} = 0$$

Note: $Z(t=0, \vec{g}) = 1$

Approach II : $\phi(x) = \sum^{-1/2}(\Lambda) \phi_0(x)$

$$\langle 0 | T \phi(x_1) \phi(x_2) \dots \phi(x_n) | 0 \rangle = G^{(n)}(x_1, \dots, x_n)$$

$$G^{(n)}(x_1, \dots, x_n) = \sum^{-n/2}(\Lambda) \underbrace{\langle 0 | T \phi_0(x_1) \dots \phi_0(x_n) | 0 \rangle}_c$$

independent of "Λ".

Computed with $H(g_1(\Lambda), g_2(\Lambda), \dots, g_n(\Lambda); m(\Lambda), \Lambda)$

$$G^{(n)} = G^{(n)}(\dots; \hat{g}_1, \hat{g}_2, \dots, \hat{g}_n(\Lambda); \hat{m}(\Lambda), t)$$

$$t = \ln \Lambda$$

$$\left\{ \frac{\partial}{\partial t} + \beta(\tilde{g}) \frac{\partial}{\partial \tilde{g}} \right\} G^{(n)} = \underbrace{\left\{ \frac{n}{2} \frac{\delta Z}{\delta \Lambda} \frac{1}{Z} \right\}}_{\gamma(\tilde{g})} G^{(n)}$$

$$\left\{ \frac{\partial}{\partial t} + \beta(\tilde{g}) \frac{\partial}{\partial \tilde{g}} + \gamma(\tilde{g}) \right\} G^{(n)} = 0$$

Note: $Z(t = \ln \frac{\Lambda_{UV}}{\Lambda_{IR}}, \tilde{g}) = 1$